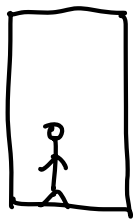


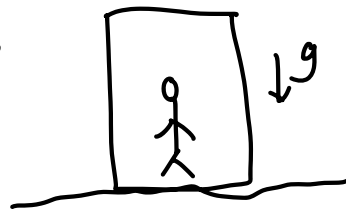
Principle of Equivalence. ①

Suppose we are in a box



We cannot tell if we are
in outer space acc. up
 $g \text{ ms}^{-2}$

or



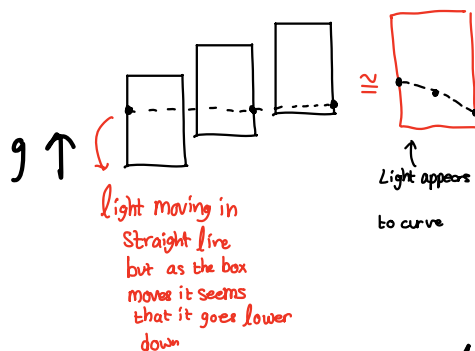
On Earth stationary
at surface
(with acc. $g \text{ ms}^{-2}$
down)



No Experiment can distinguish
between these 2.

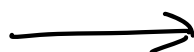
But In outer space, all points of
the box are accelerating $g \text{ ms}^{-2}$ BUT
on Earth gravity varies with height
 g at bottom is slightly greater than
top. $\rightarrow$ That difference is the
TIDAL FORCE.

② Light bends in a grav. field.



Light moving in
straight line
but as the box
moves it seems
that it goes lower
down

Light appears
to curve



But since equivalence
principle says this
is same for Earth,

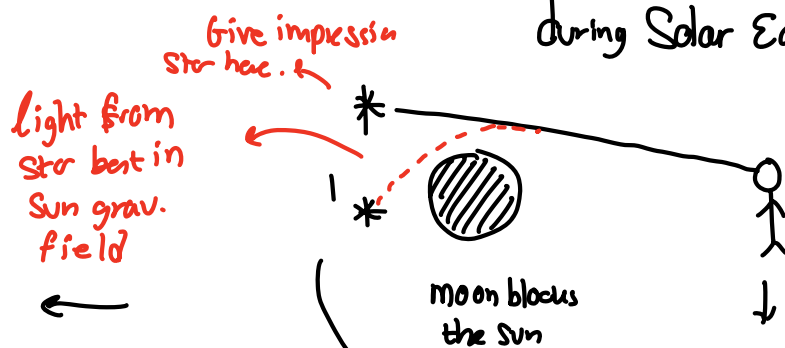
The same thing
should happen on
Earth i.e. Light
bends in a grav. field.



Shown to be true
during Solar Eclipses.

As We shine the light from the
left, the box is also moving up.
As the light reaches RHS,

Shouldn't have been
possible to see.



Give impression
star has. $\leftarrow$

Light from
star bent in
Sun grav.
field

moon blocks
the sun

But Astronomers knew
Star was located behind the
sun.

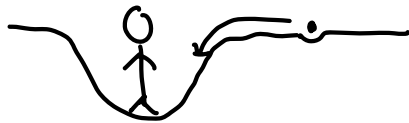
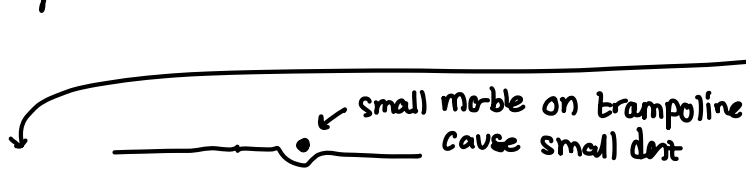
Could see
star slightly
right of sun

Problem for Einstein - Newton Grav Law

$$F = \frac{G M m}{r^2}$$

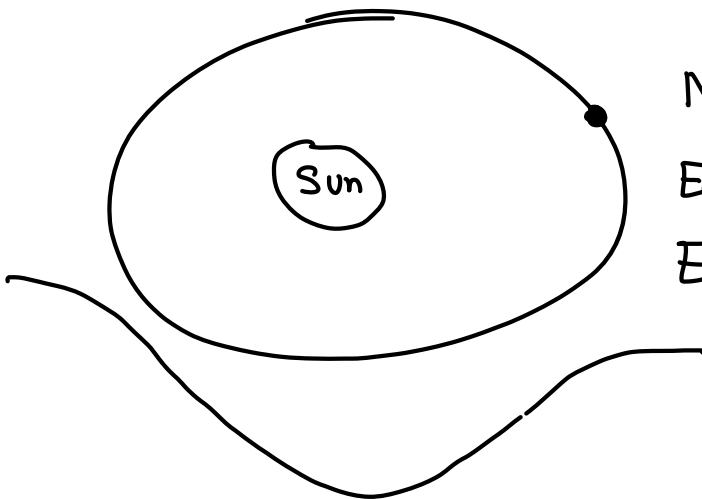
Light was bending in a field
but photons have no mass!

Instead of gravity
Einstein Propose: All forms
of motion are represented by
movement in curved space time.



Movement of marble towards larger pit
is [the shortest path in curved space time]
↳ Einstein

Newton Said - Movement towards
person is caused by grav. Attraction.

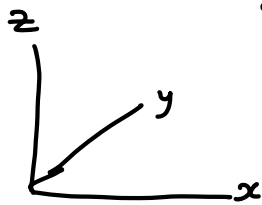


Newton: Grav. Force provides $F_{\text{centripetal}}$ to keep
Earth in Orbit.

Einstein: Massive Sun curves space time & Earth
goes round that curve in straight line in
space time.

SpaceTime - Space & Time as one set of Coordinate

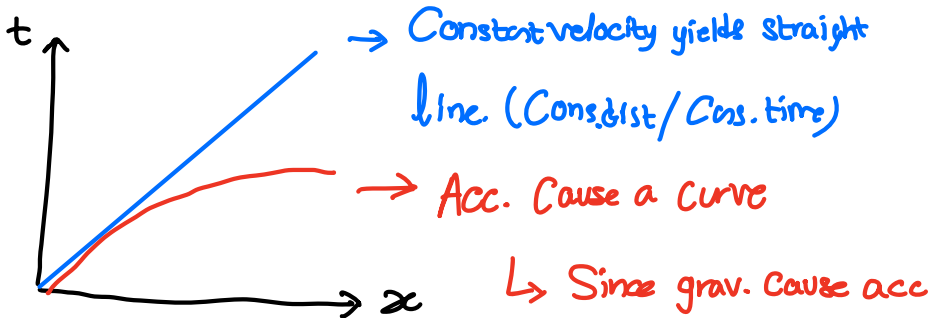
Space:



3-Orthogonal axes (3D)

+ Time = 4D Spacetime
*
↓
Difficult to draw!

* Draw 1 Space dimension + time



⇒ grav cause curve in spacetime COORDINATE.

Magnifier

Einstein thought: SpaceTime must be curved since acc.
has the effect of creating a curve in SpaceTime Coordinate.

Einstein Field Equations

Stress-Energy-Momentum Tensor.

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R + g_{\mu\nu} \Lambda = \frac{8\pi G}{c^4} T_{\mu\nu}$$

$R_{\mu\nu}$ ↑ Ricci Curvature Tensor
 $g_{\mu\nu}$ ↑ metric tensor
 R ↓ Curvature Scalar
 Λ ↑ Cosmological Constant
 $T_{\mu\nu}$ ↓ Refers to mass-energy
 $\frac{8\pi G}{c^4}$ ↓ Refers to Curvature of space-time

There are actually many eqns: indices μ & ν represent dimensions of

Spacetime: $\begin{matrix} 0 & \text{time} \\ 1 & x \\ 2 & y \\ 3 & z \end{matrix}$
Combination of μ & ν ←

⇒ 16 Variations of this eqn

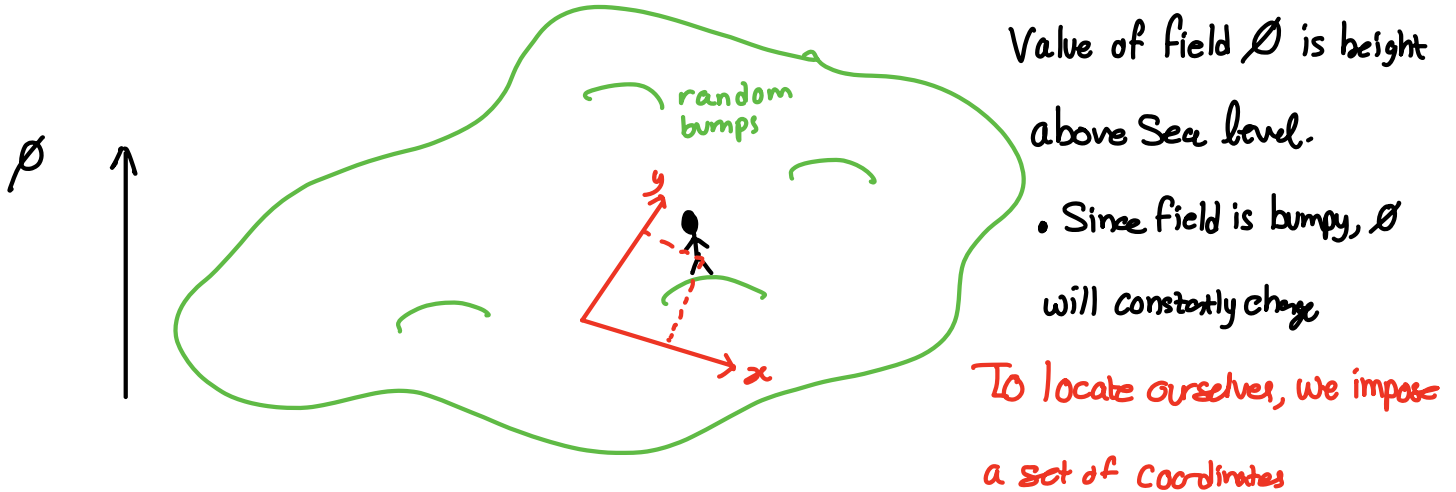
10 eqns But 6 are duplicates

Mass tell space time how to curve & Curved space-time tells mass how to move.

Metric Tensor

Consider any field : Temperature, electromagnet, gravitational, ...

↳ Land field



- Uniquely define posn and for every posn I take

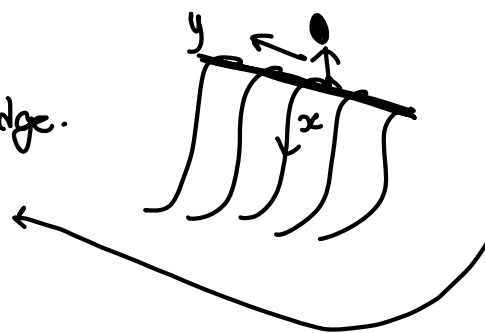
I can find a value of $\phi \rightarrow$ Height above sea level of point $(x, y) \rightarrow$ In the field

How does the height change if I move a little bit.

↳ Depend which way I move.

↳ Suppose we are at the top of a ridge.

How can I tell by how much ϕ (height) will vary as I move along?



Drop considerably

$\hat{x}$ - down ridge

y - Along ridge (flat)

↳ Stay level ground.

Consider a 1:10 gradient $\rightarrow$ Each 10 m I walk along I drop 1 m down



$$\Rightarrow d\phi = \frac{d\phi}{dx} dx$$

Change in height = Gradient $\times$ distance walked along.

$$d\phi = \frac{1}{10} \cdot 5 = \frac{1}{2} \text{ m} \rightarrow \text{If I walk 5 m along a 1:10 gradient, my height changes by } \frac{1}{2} \text{ m}$$

Change in value of field = Gradient $\times$ distance travelled.

$\hookrightarrow$ Note: Usually field is bumpy & no clear gradient is obtained.
But for very small distances ($d\phi, dx$) this is true!

But direction matters

$$d\phi_x = \frac{d\phi}{dx} dx \quad d\phi_y = \frac{d\phi}{dy} dy$$

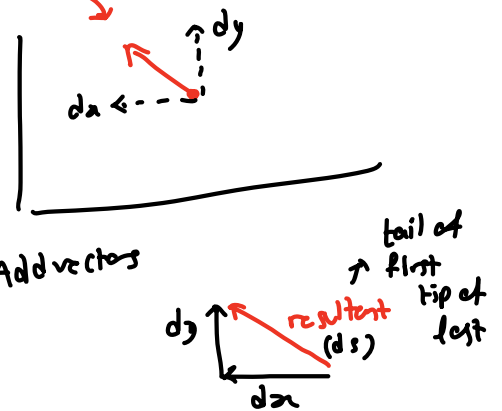
$\uparrow$ change in height while walking in x direction

$\uparrow$ gradient in y direction

Not the same!

Ridge: x is steep $\frac{d\phi}{dx}$
 y is level $\frac{d\phi}{dy}$

Suppose we walk in combination of x, y (diagonal)



Pythagoras: $ds^2 = dx^2 + dy^2$
 $\uparrow$ magnitude

Sols

vector Addition: $d\vec{s} = d\vec{x} + d\vec{y}$

Separate $\leftarrow$ Change in height:

from $d\vec{s} = d\vec{x} + d\vec{y}$
line not derived directly from this?

$$d\phi_s = d\phi_x + d\phi_y$$

Along paths

change in height along path x

for y

These are vectors

$$d\phi_s = \frac{d\phi}{dx} \cdot dx + \frac{d\phi}{dy} \cdot dy$$

$$d\phi_s = \frac{\partial \phi}{\partial x} dx + \frac{\partial \phi}{\partial y} dy$$

$\downarrow$ Partial derivative w.r.t x only as ϕ could be a function of both x & y .

No nonrelative change

$$x \rightarrow x'$$

$$y \rightarrow x^2$$

3D : $\begin{matrix} x \\ y \\ z \end{matrix}$ } usual coords. If we need to add other coords.
time relabelling works better.
 $\begin{matrix} -x^0 \\ x \\ y \\ z \end{matrix} \rightarrow \begin{matrix} -x^1 \\ -x^2 \\ -x^3 \end{matrix}$

$$d\phi_s = \frac{\partial \phi}{\partial x^1} dx^1 + \frac{\partial \phi}{\partial x^2} dx^2 + \dots$$

↑
on
drops
at we interpret
it as change in
field value

↑ If needed.

$$\Rightarrow d\phi = \sum_n \frac{\partial \phi}{\partial x^n} dx^n \quad \text{--- (1)}$$

$$n=1 \Rightarrow \frac{\partial \phi}{\partial x^1} dx^1$$

$$n=2 \Rightarrow \frac{\partial \phi}{\partial x^2} dx^2$$

Problem with GR : Any rules we come up with must be independent of
FOR.

Special Relativity : Different observers measure different distances & time depending
on their relative velocity $\rightarrow$ Length contraction & Time dilation.

$\hookrightarrow$ For something to be universally true, it must be true in all FOR.